The local and spatial form of the momentum balance equation is

$$\nabla \cdot \boldsymbol{\sigma} + \rho \, \mathbf{b} = \rho \, \frac{D \mathbf{v}}{D t}. \tag{1}$$

For review, see the lecture notes on continuum kinematics or watch and read the module "Eulerian Lagrangian Description" on

http://web.mit.edu/hml/ncfmf.html.

- This equation should always hold as long as Newton's 2nd law of motion is valid no matter what material we are dealing with.
- What distinguishes solid from fluid is the "behavior" or response to loading, which is described macroscopically by a constitutive relation.
- The boundary between solid-like and fluid-like behaviors is often blurry but one way of defining fluid is to see whether a material deforms indefinitely for given shear stress.

- If a material subjected to a fixed value of shear stress continues to deform without stopping, i.e. "flows", the material is treated as a fluid.
- Since the amount of deformation is indefinite, a proper way of measuring deformation of fluid is to measure the rate of deformation.
- The rate of deformation is quantified by the strain rate tensor:

$$\dot{\boldsymbol{\varepsilon}} = \frac{1}{2} \left[\nabla \boldsymbol{\mathsf{v}} + (\nabla \boldsymbol{\mathsf{v}})^T \right] = \frac{1}{2} \left[\frac{\partial \boldsymbol{\mathsf{v}}}{\partial \boldsymbol{\mathsf{x}}} + \left(\frac{\partial \boldsymbol{\mathsf{v}}}{\partial \boldsymbol{\mathsf{x}}} \right)^T \right], \quad (2)$$

where \mathbf{v} is the spatial velocity and x is the spatial coordinates.

In the matrix form, the strain rate tensor looks like

$$\begin{pmatrix} \frac{\partial v_1}{\partial x_1} & \frac{1}{2} \left(\frac{\partial v_1}{\partial x_2} + \frac{\partial v_2}{\partial x_1} \right) & \frac{1}{2} \left(\frac{\partial v_1}{\partial x_3} + \frac{\partial v_3}{\partial x_1} \right) \\ \frac{1}{2} \left(\frac{\partial v_1}{\partial x_2} + \frac{\partial v_2}{\partial x_1} \right) & \frac{\partial v_2}{\partial x_2} & \frac{1}{2} \left(\frac{\partial v_2}{\partial x_3} + \frac{\partial v_3}{\partial x_2} \right) \\ \frac{1}{2} \left(\frac{\partial v_3}{\partial x_1} + \frac{\partial v_1}{\partial x_3} \right) & \frac{1}{2} \left(\frac{\partial v_3}{\partial x_2} + \frac{\partial v_2}{\partial x_3} \right) & \frac{\partial v_3}{\partial x_3} \end{pmatrix}$$

In the indicial notation,

$$\dot{\varepsilon}_{ij} = \frac{1}{2} \left(\frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)$$
(3)

- Next, we are going to make two assumptions: the flow is incompressible and Newtonian.
- An **incompressible** flow has zero volumetric strain rate:

$$\dot{\boldsymbol{e}} = \operatorname{tr}(\dot{\boldsymbol{\varepsilon}}) = (\varepsilon_{11} + \varepsilon_{22} + \varepsilon_{33}) = \left(\frac{\partial v_1}{\partial x_1} + \frac{\partial v_2}{\partial x_2} + \frac{\partial v_3}{\partial x_3}\right) = 0.$$
(4)

The Cauchy stress tensor can always be decomposed into pressure and deviatoric stress terms.

$$\boldsymbol{\sigma} = -\mathbf{p}\,\mathbf{I} + \boldsymbol{\tau},\tag{5}$$

where $p = -tr(\sigma)/dim$ and the minus sign implies that compressional pressure is positive.

 Likewise, the strain rate tensor can be decomposed into volumetric and deviatoric components.

$$\dot{\boldsymbol{\varepsilon}} = \dot{\boldsymbol{\theta}} + \dot{\boldsymbol{\epsilon}},$$
 (6)

A Newtonian flow has linear relationship between deviatoric strain rate and stress:

$$\boldsymbol{\tau} = 2\mu \dot{\boldsymbol{\epsilon}} \tag{7}$$

where μ is called (Newtonian) viscosity.

 The full constitutive relation for the incompressible Newtonian fluid becomes

$$\sigma = -p\mathbf{I} + \tau$$

= -p\mathbf{I} + 2\mu\dot{\epsilon} (8)

In the indicial notation,

$$\sigma_{ij} = -\mathbf{p}\,\delta_{ij} + \mu\left(\frac{\partial \mathbf{v}_i}{\partial \mathbf{x}_j} + \frac{\partial \mathbf{v}_j}{\partial \mathbf{x}_i}\right) \tag{9}$$

- Note that zero volumetric strain rate does not mean zero pressure in incompressible fluids.
- Let's rewrite the momentum balance equation (1) in terms of pressure and deviatoric stress to get the Navier-Stokes equation:

$$\rho \frac{D\mathbf{v}}{Dt} = \nabla \cdot \boldsymbol{\sigma} + \rho \, \mathbf{b} = \nabla \cdot (-\rho \, \mathbf{I} + \boldsymbol{\tau}) + \rho \, \mathbf{b}$$

$$= -\nabla \rho + \nabla \cdot \boldsymbol{\tau} + \rho \, \mathbf{b}.$$
(10)

For the Newtonian fluid,

$$\rho \, \frac{D\mathbf{v}}{Dt} = -\nabla \rho + \nabla \cdot (2\mu\dot{\boldsymbol{\epsilon}}) + \rho \, \mathbf{b}. \tag{11}$$

The notorious difficulty with solving the Navier-Stokes equation partly originates from the non-linear advection term involved in the material derivative:

$$\frac{D\mathbf{v}}{Dt} = \frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v}.$$
 (12)

► However, for steady-state flows (∂v/∂t = 0) with very low speed and velocity gradient, the whole left hand side of (11) can be ignored. Creeping flows of rocks satisfy this condition.

So, the Navier-Stokes equation for the steady (∂v/∂t = 0), smooth (i.e., |∇v| ≪ 1) and slow (|v| ≪ 1) Newtonian fluid becomes

$$\mathbf{0} = -\nabla \boldsymbol{\rho} + \nabla \cdot (\mathbf{2}\mu \dot{\boldsymbol{\epsilon}}) + \rho \, \mathbf{b}. \tag{13}$$

In the indicial notation,

$$\mathbf{0} = -\mathbf{p}_{,i} + \left[\mu\left(\frac{\partial \mathbf{v}_i}{\partial \mathbf{x}_j} + \frac{\partial \mathbf{v}_j}{\partial \mathbf{x}_i}\right)\right]_{,j} + \rho \mathbf{b}_i. \tag{14}$$

where (), *i* denotes partial derivative with respect to x_i with *i*=1,2,3, and the indicial notation of strain rate (3) is also used.

lf viscosity (μ) is constant,

$$\mathbf{0} = -\mathbf{p}_{,i} + \mu \left(\frac{\partial \mathbf{v}_i}{\partial \mathbf{x}_j} + \frac{\partial \mathbf{v}_j}{\partial \mathbf{x}_i} \right)_{,j} + \rho \mathbf{b}_i. \tag{15}$$

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The pressure *p* can be decomposed into hydrostatic pressure (e.g., ρgx₂) and dynamic pressure (*P*):

$$\boldsymbol{p} = \rho \boldsymbol{g} \boldsymbol{x}_2 + \boldsymbol{P}. \tag{16}$$

If we further assume that the body force is solely due to gravity, then the gradient of the hydrostatic pressure and the body force terms cancel out each other.

$$-\frac{\partial \rho g x_2}{\partial x_2} + \rho g = 0.$$
 (17)

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Then, we have only dynamic pressure term left:

$$\mathbf{0} = -\mathbf{P}_{,i} + \mu \left(\frac{\partial \mathbf{v}_i}{\partial \mathbf{x}_j} + \frac{\partial \mathbf{v}_j}{\partial \mathbf{x}_i} \right)_{,j}.$$
 (18)

In a 2D case, denoting x₁ and x₂ as x and y, v₁ and v₂ as u and v, respectively, we can write the above equation as follows:

for *i* = 1:

$$0 = -\frac{\partial P}{\partial x} + \mu \left(\frac{\partial}{\partial x} \left(\frac{\partial u}{\partial x} + \frac{\partial u}{\partial x} \right) + \frac{\partial}{\partial y} \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) \right)$$

$$= -\frac{\partial P}{\partial x} + 2\mu \frac{\partial^2 u}{\partial x^2} + \mu \left(\frac{\partial^2 u}{\partial y^2} + \frac{\partial^2 v}{\partial x \partial y} \right).$$
 (19)

for i = 2:

$$0 = -\frac{\partial P}{\partial y} + \mu \left(\frac{\partial}{\partial x} \left(\frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} \right) + \frac{\partial}{\partial y} \left(\frac{\partial v}{\partial y} + \frac{\partial v}{\partial y} \right) \right)$$

$$= -\frac{\partial P}{\partial y} + \mu \left(\frac{\partial^2 v}{\partial x^2} + \frac{\partial^2 u}{\partial x \partial y} \right) + 2\mu \frac{\partial^2 v}{\partial y^2}.$$
 (20)

The incompressibility condition for 2D becomes

$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0 \tag{21}$$

By differentiating this with respect to x, we get

$$\frac{\partial^2 \mathbf{v}}{\partial \mathbf{x} \partial \mathbf{y}} = -\frac{\partial^2 u}{\partial \mathbf{x}^2}.$$
 (22)

Likewise, with respect to y,

$$\frac{\partial^2 u}{\partial x \partial y} = -\frac{\partial^2 v}{\partial y^2}.$$
 (23)

Simplifying the 2D equation with these identities, we get

$$0 = -\frac{\partial P}{\partial x} + \mu \left(\frac{\partial^2 u}{\partial x^2} + \frac{\partial^2 u}{\partial y^2} \right)$$

$$0 = -\frac{\partial P}{\partial y} + \mu \left(\frac{\partial^2 v}{\partial x^2} + \frac{\partial^2 v}{\partial y^2} \right).$$
(24)

Equations in (24) are the governing equations for the motion of a steady, smooth, slow, Newtonian fluid in 2D and the same with (6-67) and (6-68) in T&S.

In 1D,

$$0 = -\frac{dP}{dx} + \mu \frac{d^2 u}{dx^2}.$$
 (25)

This is the same with (6-10) in T&S.

The first half of the fluid mechanics chapter is just a collection of applications of these equations. Note that 1D equation is an ODE, which can be solved very easily.